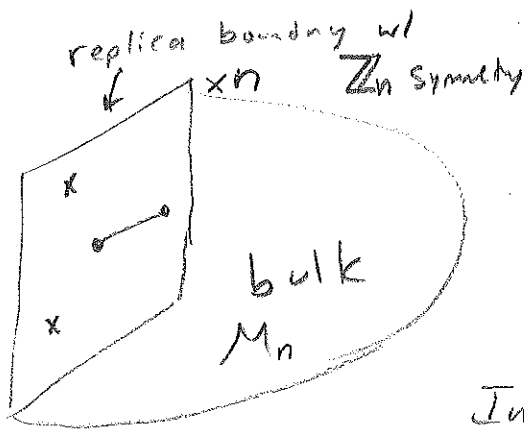


# PROOF FROM GRAVITATIONAL PATH INTEGRAL

3.5



$$Z_{\text{boundary}} = Z_{\text{bulk}}$$

$$I_{\text{eff}}(\text{CFT}) = I_{\text{GR}}(\text{bulk})$$

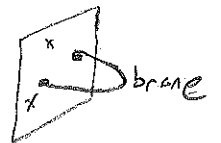
Conical singularity on boundary

might think there should be a singularity in the bulk, but this would violate Einstein's

Instead, find smooth bulk instanton that dominates saddle point w/ specified b.c.

Assume  $Z_n$  symmetry not spontaneously broken

Orbitoid  $O_n = M_n / Z_n$  has original  $\partial M_n$  as boundary now has conical singularity — "brane" w/ tension solves  $I = I_{\text{GR}} + I_{\text{brane}}$



NB this brane is entirely fictitious — does not really contribute to  $Z$ .  $I_{\text{brane}} \propto (n-1)A/G$

Define  $I[O_n] = I[M_n] / n$  by stipulation!

[MORE EQ]

now can analytically continue to  $n \rightarrow 1$  & it's well defined, [MAKE REF OF  $M_n$  MORE EXPLICIT]

Note,  $n=1$  is original solution to EOM, so  $\partial_n (I_{\text{GR}} + I_{\text{brane}})|_{O_n} = 0$

We can switch from  $I_{\text{GR}} \rightarrow -I_{\text{brane}} \propto (n-1) \frac{A}{G} [O_n]$

[DONE BETTER ON NEXT PAGE]

use  $S_n = \frac{1}{1-n} \ln \frac{Z_n}{Z_1^n}$

[EXTENSION TO COVARIANT CASE]

[Should this be post Wald S?]  
[order of limits issue]

[MENTION CASINI-HUERTA-MYERS DIO STATIONARY VERSION OF HYPERBOLIC BH]



I[O\_n] = 1/n I[M\_n] @ integer n

this now defines I[M\_n] @ noninteger n

n=1 solves EOM for S\_GR

action of conical deficit

so d\_n(I[O\_n] - (1-n) A[O\_n]/G) |\_{n=1} = 0

I\_deficit = -I\_brane

if solutions to

n = 1 + epsilon

I\_bulk\_smooth + I\_def + I\_brane

so I(O\_{1+epsilon}) - epsilon A[O\_{n=1}]/G = I(O\_1) (= I(M\_1))

by varying the area along the brane

[FIXED SIGNS]

I(M\_{1+epsilon}) = (1+epsilon) I(O\_{1+epsilon}) = (1+epsilon) I(M\_1) + epsilon A[ext]/G

S = + 1/epsilon (I(M\_{1+epsilon}) - (1+epsilon) I(M\_1)) = A[ext]/G

[NOT KEEPING TRACK OF NUMERICAL FACTORS...]

- from 1/(1-n)
- from I = -ln z

earlier Fursaev Proof erroneous as pointed out by Headrick.

FLM?

CHECKING SIGNS

+R -> +ln Z -> -I
I = -R / (16\*pi\*G)

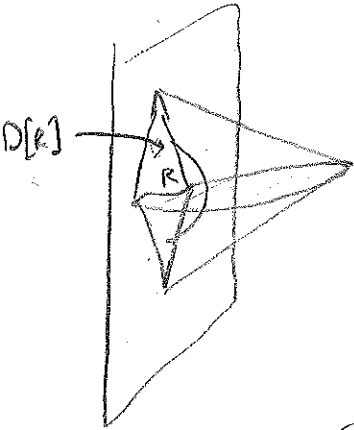
Did conical variable in bulk, but result not saddle point, & gets wrong result for Ray's, (but right for von Neumann.)

# BULK RECONSTRUCTION

3.7

Suppose we have region  $R$  on the boundary

Armed only with  $P_R$ , how much of the bulk can we reconstruct?



$$\text{Causal wedge} = I^-(D[R]) \cap I^+(D[R])$$

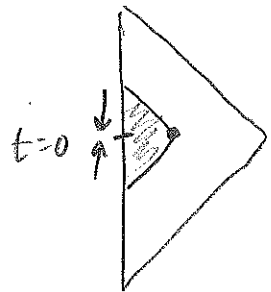
region for which signals can leave  $D[R]$  & return could be probed thru experiment

HKLL — nonstandard Cauchy problem

Fourier in transverse directions

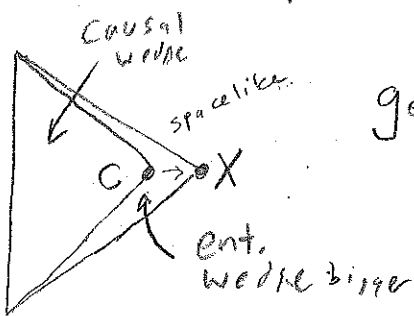
reduce to  $|t|$  Cauchy evolution

evolve back to  $t=0$  using CFT evolution



BUT GOOD REASONS TO THINK YOU CAN GO FARTHER

[OPEN QUESTION IF NO SYM HOW TO GENERALIZE]



generically,  $X$  lies deeper in bulk than  $C$

(concedes in special cases like 1 interval AdS)

$$A[C] \geq A[X]$$

= whenever analytically known, but

$A[X]$  is special, should be at boundary of special region

coarse-grained entropy???

fine grained  $S_{CFT}$

$A[C]$  fails SSA,  $S[R] = S[\bar{R}]$  etc.

natural conjecture, can reconstruct all fields in ent wedge.

evidence: consistency conditions

$$JLMS \text{ mod } K_{CFT} = K_{\text{bulk}} + \frac{A[\text{ext}]}{4Gh} + \text{DHW}$$

tensor networks: HPPY model & error correction (ADH)

QUANTUM ERROR CORRECTION?



if time...

# HORIZONS

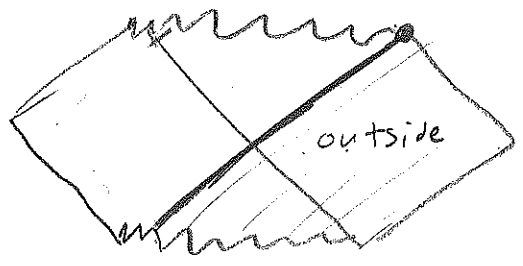
"future  
causal horizon":

boundary of past of locus of points @ fut infinity.

Examples:

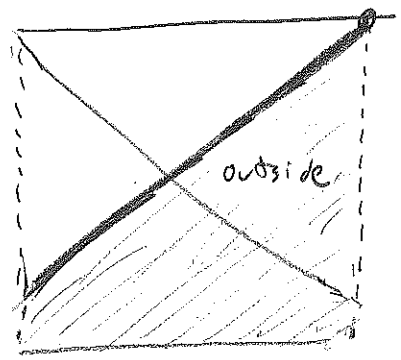
"past CH" = time reverse

BH (+ pert.)



location is still defined  
after perturbations

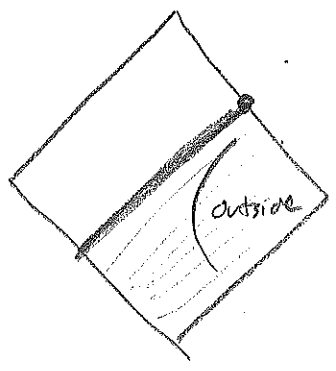
dS (+ pert.)



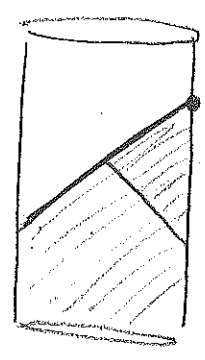
"teleological  
boundary  
condition"

Rindler  $R_{1,1} \times R_{D-2}$

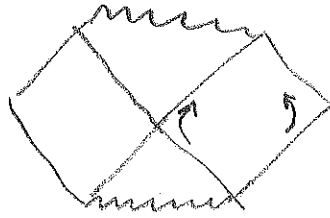
[DISCUSS APPARENT  
HORIZONS?]



Rindler AdS



# Black Hole Thermodynamics



Energy  $E = M$

Temp  $T = \frac{\hbar \kappa}{2\pi}$

where  $\kappa =$  surface gravity =

ratio of  $\frac{\text{boost @ horizon}}{\text{time @ } \infty}$  for Killing horizon

Entropy  $S = \frac{A}{4G\hbar} \quad (c = k_B = 1)$

## ZEROth LAW

T constant @ equilibrium



$\kappa$  constant for stationary Killing horizon

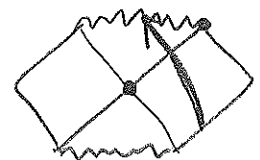
follows from "dominant energy condition"  $T_{ab} t^a t^b \geq 0$   
 or from sufficient symmetry (static, or axisymmetric & "t- $\phi$ " reflection symmetric)

future timelike



## ENERGY CONSERVATION

E is encoded in ADM mass @  $\infty$ ,  
 which is conserved due to diffeo symmetry



## "FIRST LAW" / CLAUSIUS RELATION

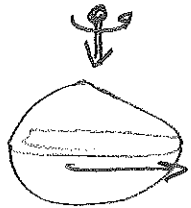
$dE = T dS [ + \underbrace{\Omega}_{\text{chemical potentials for angular momentum } J} dJ + \underbrace{\phi}_{\text{charge } Q} dQ + \dots ]$  (for linear path to Killing horizon from beginning to end)

Can transform into co-rotating,  $\phi = 0$  frame to get  $dE = T dS$

# THIRD LAW

(4.3)

cannot by any finite process reach "extremal" BH  
w/  $T=0$  (e.g. from large enough  $J, Q$ )



Can't force spin in w/o  
adding energy

Does NOT imply  
 $S=0$  when  $T=0$ !

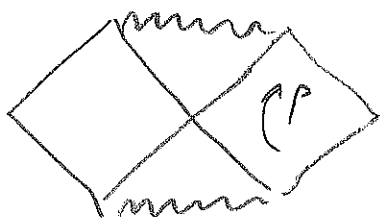
# SECOND LAW

Classically,  $A$  is increasing if null energy condition

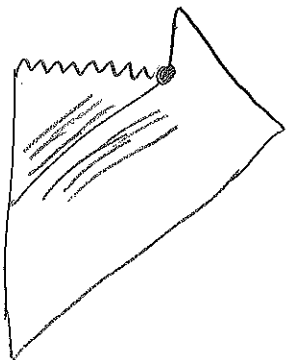
$$T_{ab} k^a k^b \geq 0$$

↑ null vector

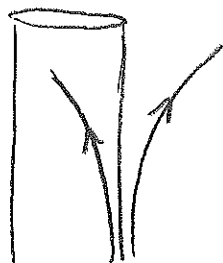
Take QM into account, BH's radiate...



$P$  is thermal restricted to  
exterior in HH state.



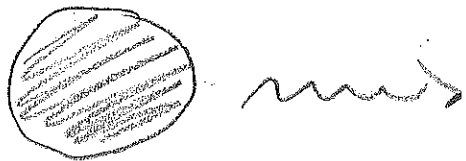
in case of BH w/ forms from  
collapse, follows from Unruh effect  
of short distance modes at early times.



So a BH in Unruh state (no incoming rad)  
will shrink with time, eventually reach  
Planck size (and probably disappear).

# GENERALIZED ENTROPY

4.4



BH can radiate to exterior; so what matters is sum of BH ent. and exterior entropy  $S_{out}$

$$S_{gen} = \frac{A}{4G\hbar} + S_{out} + \text{higher-curvature corrections}$$

GSL:

$$\left. \frac{dS_{gen}}{dt} \right|_{\text{causal horizon}} \geq 0$$

Quanta/Thermal

$-\text{tr}(\rho \ln \rho)$ , divergent.

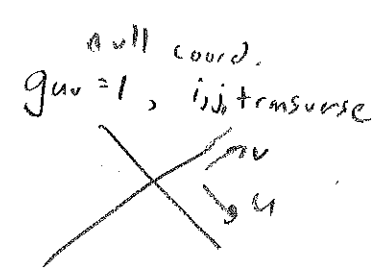
need regulator  $\epsilon$ , divergences absorbed into  $\frac{1}{G}$ , counterterms.

Higher Curv./Stringy

$$I = \int d^D x \sqrt{g} L, \quad L = \frac{R}{16\pi G} + \alpha(R^2 \dots)$$

"Conical"  $\Downarrow$  Corrections to BH entropy "replica"

$$S_{BH} = -\frac{2\pi}{\hbar} \int d^{D-2} x \sqrt{g_2} \left[ 4 \frac{\partial L}{\partial R_{\mu\nu\rho\sigma}} + 16 \frac{\partial^2 L}{\partial R_{\mu\nu\rho\sigma} \partial R_{\kappa\lambda\epsilon\zeta}} K_{ij}^{(\mu)} K_{kl}^{(\nu)} + \mathcal{O}(K^4) + \dots \right]$$



Wald, valid stationary,  $f(R)$

Solodukhin, FPS, Dong, Camps, Miaou...

$$= \frac{2\pi}{\hbar} \int \sqrt{g} \frac{\partial I}{\partial R_{abcd}} E_{ab} E_{cd}$$

where  $E = \beta$  on horizon

"Noether charge"

agrees w/ Wald 2nd Law for Killing + 1st order pert. for all higher curv. thys.

quadratic

$f(\text{Riemann})$   
Holographic EE

"splitting problem" derivatives of R